# VARIABILITY IN THE F1 AND F2 REGION ELECTRON DENSITY DURINGA GEOMAGNETIC ACTIVITY AT MID-LATITUDE\*

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### Abstract

This paper investigates the variation in the F1 and F2 region electron density during a geomagnetic activity, at East Asian mid-latitude stations. In this analysis, we have employed D(foF1) and D(foF2) representing deviation of the critical frequency for the F1 and F2 ionospheric regions respectively. The F1 region appears to be much more stable than the F2 layer during the storm event, as there is no significant effect on the F1 layer in most of the ionospheric stations under investigation. It was also observed that independent of the sign of the storm effect on NmF2, the electron density, if any, in the F1 region is always negative. Moreover, no F1 ionospheric response was observed at midnight (0000UT) throughout the storm event in all the stations; but recorded its maximum effects between 0600UT-1800UT during the day. Conclusively, there is a considerable intra-hour variability of F2 electron density NmF2 during ionospheric disturbances.

Key words: Electron density, geomagnetic storm, critical frequency, ionospheric response.

### 1. Introduction

The geomagnetic storm is the most important phenomenon in the complex chain of solar terrestrial relations and space weather. However, the storm is supplied by the solar wind energy, captured by the magnetospheric and transffered and dissipated in the high and mid-latitude ionosphere and atmosphere (Buresova and Lastovick, 2001). According to Danilov (2001), the response of the ionospheric disturbances is different from that of the lower ionosphere. The difference is due to the differences in physical mechanisms responsible for the changes of the electron concentration [e]. While in the lower regions, the primary reason of the [e] changes is the variation of the ionization rate because of corpuscular intrusions, there is no considerable change of the ionizing source intensity in the F2 region during geomagnetic disturbances and so the electron concentration variations are due to indirect factors. The geomagnetic storm effects on the F2-region ionization at a given location depend, in a complicated way, on the time, season and storm onset time (Kilifurska, 1998; Prolss, 1993a). Physical causes of the negative and positive response of the F2- region to a geomagnetic storm have been studied exclusively (Prolss, 1993a,b; Condreseu et al., 1997; Mikhailov and Schlegel, 1998) and need not be discussed here. Our interest in this paper deals with the variation in and F1-layer ionization response to a geomagnetic storm at mid-latitude in comparison with

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changes in the F2-layer electron density NmF2 for the same storm event.

#### 2. Data and Method

The ionospheric data used in this study consists of hourly values of the F layer critical frequency foF2 obtained from some of the National Geophysical Data Center's SPIDR (Space Physics Interactive Data Research) a network of ionosonde stations located in the East Asian sector of the world: The F layer critical frequency foF2 is used because of its direct relationship with the F layer peak electron density NmF2 (which is a measure of positive or negative storm effects through its significant increases or decreases about the mean position respectively). i.e.

$$foF2 (Hz) = 9.0 x \sqrt{[NmF2] (m^{-3})}$$
 (1)

The present study is concerned with variability in the F1 and F2 region electron density during the intense geomagnetic storm of January 10-11, 1976 at midlatitude. However, the F2 region response to a geomagnetic storm is most conveniently described in terms of the normalized deviations of the critical frequency foF2 from the reference, D(foF2) (Chukwuma 2003b), where

D(foF2) = [foF2 - (foF2)ave] / (foF2)ave (2)

(2)

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Hence the data under analysis consists of D(foF2) & D(foF1) of respective hourly values of foF2 and () negative phase again through 2000UT, January 12. foF1 respectively on January 5-12, 1976. The reference for each hour is the average value of foF2 and foF1 for that hour calculated from the five quiet days in January 5-9, 1976, preceding the storm. The use of D(foF2) and D(foF1), the normalized deviations of the critical frequency rather than the critical frequency itself provides a first-order correction for temporal, seasonal and solar cycle variations, so that geomagnetic storm effects are

3. Ionospheric Observations and Results Fig. 1 and 2 shows D(foF2) and D(foF2) plots for the period of January 10-12, 1976 for the six East Asian mid-latitude ionosonde stations of Khabarovsk (48.5°N), Wakkanai (45.4°N), Akita (39.7°N), Kokunbunji (35.7°N), Yamagawa (31.2°N) and Okinawa (26.3°N). Ionospheric F region electron density is determined mainly by photoionisation, neutral composition and winds during geomagnetic quiet periods. However, in this analysis, our main interest shall be in explaining the response of the F1 and F2 regions of the ionosphere to the intense geomagnetic storm of January 10-11, 1976 and the

### F2 Region Response

The D(foF2) plot at Khabarovsk (Figure 1) shows a positive storm between 0000UT and 0300UT on January 10. Observe the D(foF2) variations show the ionosphere developing a negative storm at o the reference. It should also be observed that the peak depletion in foF2 at 0800UT, 1500UT, January 10 and 0000UT and 1000UT, January 11 coincides with the large increases in proton number density (Figuge 3b) at this same points, which according to Strickland et al (2001) indicate the arrival of a shock in the interplanetary medium. However, between 0600UT and 0900UT, January 11, a positive storm was observed, which thereafter depletes, reaching a 50% depletion level on January 11, and then begins rotating northward through January 12. Between 0600UT and 2100UT, January 12, a positive phase storm with 30% enhancement level was observed. storm through 0300UT, January 12. However, a brief positive storm was observed between 1900UT and 2100UT, January 10 and 0600UT, January 11. Furthermore, negative storm observed at this station also preceeded the intense magnetic storm. Note the 80% ionospheric response enhancement at 2300UT,

January 12, before a gradual decrease resulting in The D(foF2) plot at Akita shows a predominantly negative phase storm between 0000UT and 1000UT January 10. However, the foF2 pattern observed at this station is irregular showing a 50% ratio apiece for both positive and negative phase storms. Note the enhancement between 0000UT and 0900UT, January 11, having a 37% peak enhancement value. Available foF2 data at Kokubunji is similar to the D(foF2) plot at Akita except that a negative phase better identified (Chukwuma, 2003b). Isob norisola noristorm was observable between 0000UT and 2300UT, norisola gininessitation (1561) discussion fanuary 10. Thereafter, a positive phase storm was imminent up till pre noon hours of January 11, when it experiences a southward rotation resulting in negative phase storm which lasted till 1600UT before another enhancement to a peak value of 59% was observed. It thereafter begins to recover and maintains a negative phase storm between 0300UT and 2300UT, January 12.

The plot of the ionospheric response at Yamagawa as seen from the figure, indicates a rather irregular pattern between 0000UT and 1800UT, January 10, but more of negative storm. Note the enhancement between 0000UT and 2100UT, January 11. The irregular pattern thereafter continues through January electron density variations between the two. " zogmaiol2. it should also be noted that the 31% peak depletion dT the bedraft of the same storn event at the same storn event by an increase in proton density about the same time lor Swhich also coincides with the minimum peak value me of Dst, thus indicating the presence of an intense sm high and mid-latitude ionosphere and atmospheres 0800UT and attaining a 28% depletion level from to The situation at Okinawa is not different, an existing (8) positive phase storm was observed between 0500UT, so January 10 and 0200UT, January 11. Note thesis depletion in foF2 observed between 0200UT and of 2200UT, January 11; and 0300UT and 1200UT, mi January 12. Apart from these two points, the F2 10 response is predominantly positive. (Taming off, and predominantly positive.) reliation of the ionization rate of second star noise and the noise rate The Floregion electron density response at in Khabarovsk is shown in Fig. 2. There is no data available for January 10. However, the plot for January 11-12 shows little or no response as the observations are close to the reference line. The situation was the same at Wakkanai. Here, the The ionospheric response at Wakkanai shows that to negative storm effect observed is not up to 10% there were no immediate response between 0000LJT depletion and so cannot be regarded as an appreciable and 0700UT, January 10. With effect from 0700UT, m response. The D(foF1) plot at Akita showed no di January 10, the ionosphere recorded a negative phase in immediate response in the ionosphere until around of 0300UT, January 11 showing a rather weak storm effect up till 2300UT. However, with effect from M 0300UT, January 12, it begins to experience negative storm of about 23% depletion level till 2300UT of geomagnetic storm at mid-latitud 57:000/01:24 same day. The situation at Kokubunji (Fig. 2) shows an (Hionospheric activity beginning around 0800UT, +

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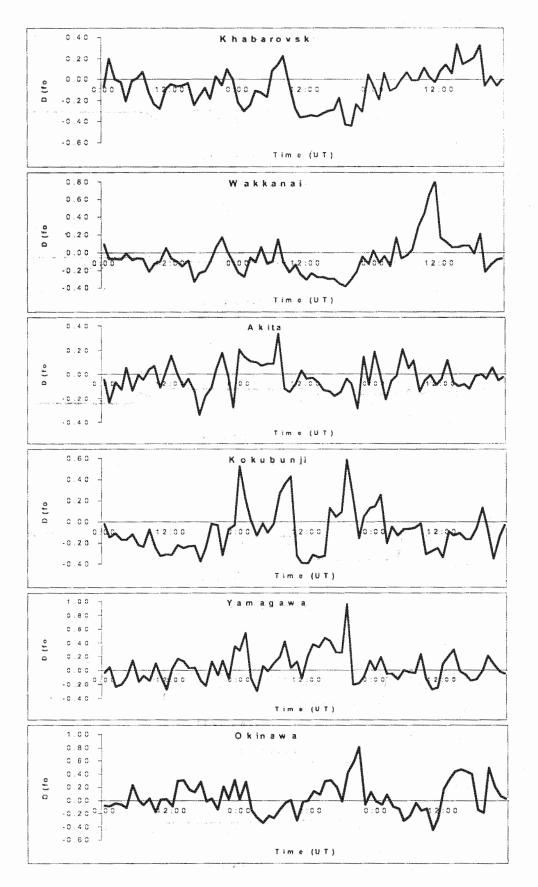
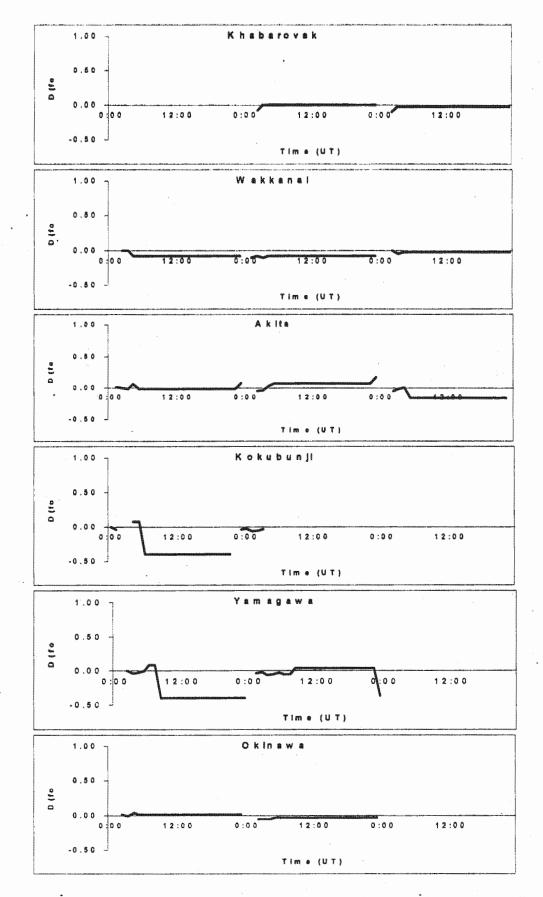
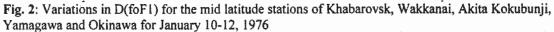


Fig. 1: Variations in D(foF2) for the mid latitude stations of Khabarovsk, Wakkanai, Akita Kokubunji, Yamagawa and Okinawa for January 10-12, 1976





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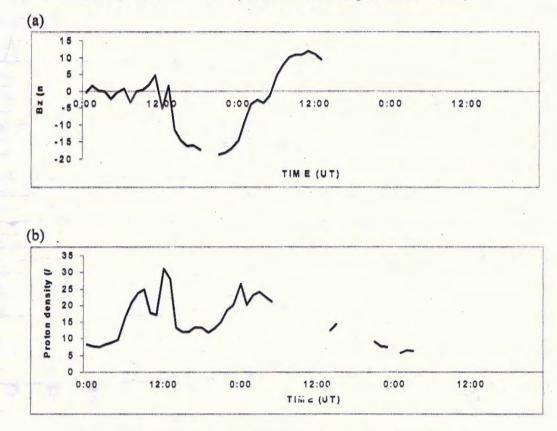


Fig. 3: Composition of interplanetary observations for January 10-12, 1976,

to about 40% depletion, and thereafter maintaining this value till 2000UT of same day. Note that the time of activity increase at this station coincides with the time of storm sudden commencement (SSC) on the Bz plot (Fig. 3a). Paucity of data would not allow comment on the remaining days. The ionospheric response at Yamagawa is similar to that of Kokubunji except for the negative phase storm experienced around 2300UT, January 11. The condition at Okinawa could also be regarded as no effect on the F1 region.

## 4. Discussion

It is well established that the Bz component of the IMF is the most important influence on the magnetosphere and high and mid-latitude ionosphere as it controls the fraction of the energy in the solar wind which is extracted by the magnetosphere. Hence, the storm experienced at some of the mid latitude stations after storm commencement appear to be caused by the short duration southward turning of Bz giving  $\delta Bz = -12nT$  between 0600UT and 0800UT on January 10. It thus appear that this southward turning with  $\delta Bz = -12nT$  may have been accompanied by an increase in solar wind dynamic pressure which led to an enhanced coupling between the solar wind and the terrestrial magnetosphere that significantly increased the geoeffectiveness of the solar wind. (Chukwuma, 2007)

The appearance of the positive storms at the high latitude stations under investigation is as a result of energy being injected into the polar upper atmosphere as the solar wind become geoeffective; which in turn launches a traveling atmospheric disturbance (TAD) which propagates with high velocity (Danilov, 2001). This TAD carries along equatorward- directed winds of moderate magnitude. At high latitudes, these meridional winds drive ionization up inclined magnetic field lines and cause uplifting of the F layer, leading to an increase in the ionization density i.e. positive storm.

Table 1: Types of Atoms/Molecules and Maximum Number of Electrons in each Region.

Regions	Atom/Molecule	Maximum no of electron
D	N <sub>2</sub> , O <sub>2</sub>	$10^{8} - 10^{10}$ electron/m <sup>3</sup> (day)
E	$N_2 > O_2 > O$	10 <sup>11</sup> e/m <sup>3</sup> (day), 10 <sup>9</sup> e/m <sup>3</sup> (night)
F	$N_2 > O > O_2$	$10^{12} e/m^3$ (day)
F <sub>2</sub>	$0 > N_2 > O_2$	1012 e/m3 (day), 5x1010 e/m3 (night)

(after Oyinloye, 1988)

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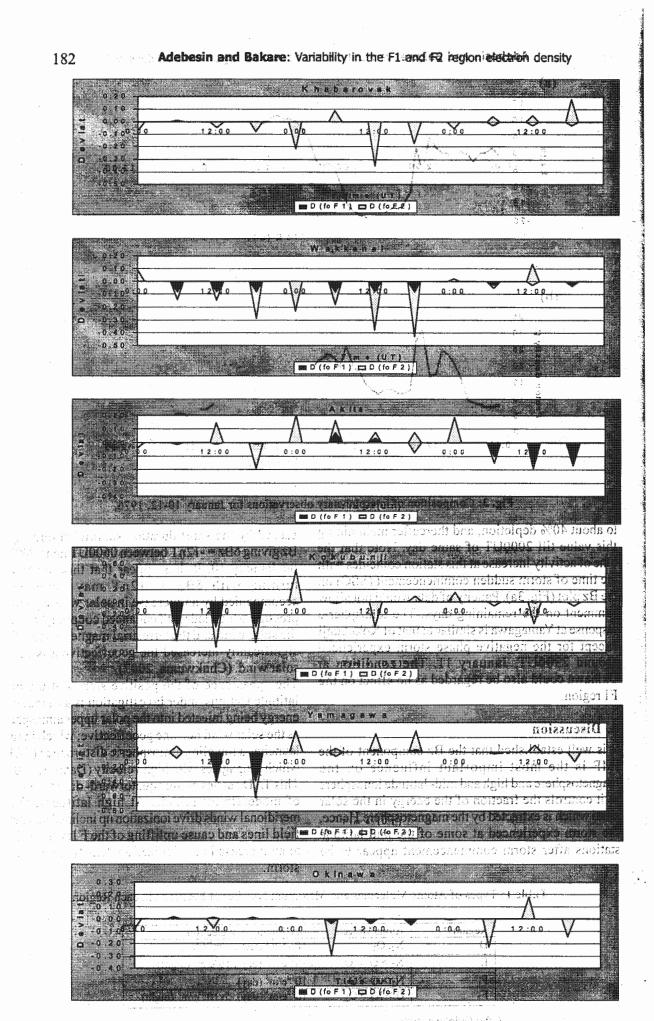


Fig. 4: Six hours interval sampling in the variations in D(foF1) and D(foF2) for the mid latitude stations for January 10-12, 1976.

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The observed decrease in foF2 during the storm is related to the neutral composition disturbances. Heating at auroral and high latitudes causes expansion of the neutral atmosphere, and enhanced neutral winds carry disturbed composition. However, enhancement in the mean molecular mass in the neutral composition disturbance zone leads to an increase in the loss rate of ions, resulting in a decrease of the ionospheric plasma density and thus a negative storm. Strickland et al (2001), had shown that negative ionospheric storm effects are indeed correlated with the region of enhanced molecular mass.

However, for a clearer picture, a six-hour interval plot was made for both D(foF2) and D(foF1) on same panel to investigate the variability between the two (Fig. 4). From the figure, the following characteristics were observed.

i) the storm event does not have any effect on the F1 region electron density at midnight (0000UT) in all the stations.

ii) the maximum effects on the F1 ionosphere is observed between 0600UT-1800UT during the daytime.

iii) Independent of the sign of the storm effect on NmF2, the effect on electron density, if any, in the F1 region has always been negative.

iv) the F1 region appears to be much more stable than the F2 layer during the geomagnetic activity as there is no significant effect on the F1 layer in most of the stations under investigation.

v) there is a considerable intra-hour variability of NmF2 during the event.

The above observed characteristics suggest the fact that when we go down from the F2 region maximum, the effect of geomagnetic storms on the neutral thermosphere becomes less dramatic. According to Buresova and Lastivicka (2001), the influence of ionization and photo-chemistry processes on the ionospheric storm becomes more important due to shorter lifetime of free electrons in the more dense atmosphere. Thus, the F1 region is the region where both the changes in the neutral neutral atmosphere (dominant in F2 region) and the changes in the ionization rate and photo-chemistry (dominant in the lower ionosphere) play an important role. [Table 1 shows types of atoms/molecules and maximum number of electrons in each region].

#### 5. Conclusion and Suggestion

The variability of F1 and F2 region response to the geomagnetic activity of January 10-11, 1976 at East Asian mid-latitude stations have been studied; and it was observed that the F1 region appears to be much more stable than the F2 layer during the storm event, as there is no significant effect on the F1 layer in

most of the ionospheric stations under investigation. It was also observed that independent of the sign of the storm effect on NmF2, the electron density, if any, in the F1 region is always negative. Moreover, no F1 ionospheric response was observed at midnight (0000UT) throughout the storm event in all the stations; but recorded its maximum effects between 0600UT-1800UT during the day. Conclusively, there is a considerable intra-hour variability of NmF2 during ionospheric disturbances.

All the six mid-latitude ionospheric stations investigated are consistent with the conclusions above. In other words, the conclusions are valid for all the six stations. However, the analysis of the variability between the two layers cannot be completely concluded due to a limited data set (of considering its response to only one storm event). In light of this, we are compiling a larger database of storm events and to also investigate the phenomenon beyond the mid-latitude alone, but rather extending our investigation to high and low latitudes.

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